

To Grid or Not To Grid

Atomic Methods for Sparse Inverse Problems

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under the direction of
Prof. Martin Vetterli

co-supervision of
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Dr. Matthieu Simeoni



November 7th, 2025



Part I:

What is it about?

1. Linear Inverse Problems
2. Sparse Signals
3. Penalized Optimization
4. Atomic Reconstruction Methods



“Méthodes atomiques
pour la résolution de
problèmes inverses
parcimonieux”



“En science, un **problème inverse** est une situation dans laquelle on tente de déterminer les causes d'un phénomène à partir des observations expérimentales de ses effets.”

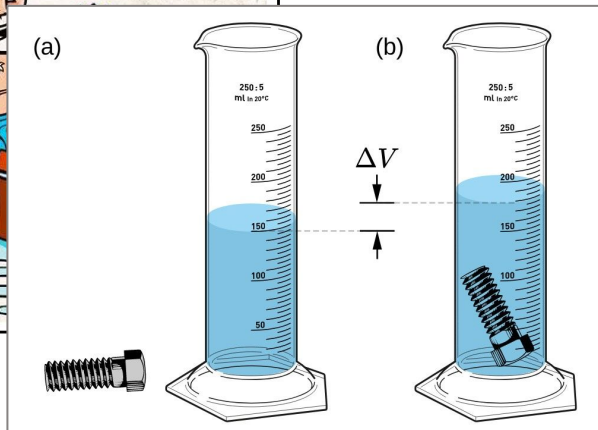
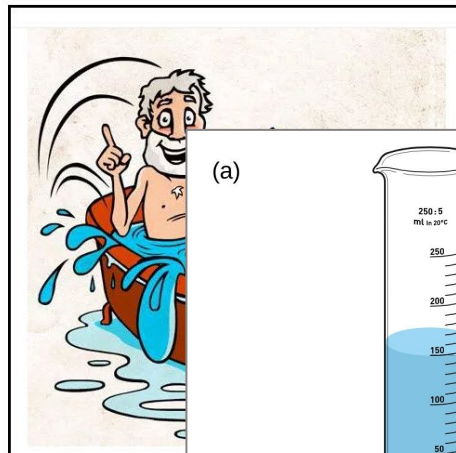
Wikipedia, FR

“An **inverse problem** in science is the process of calculating from a set of observations the causal factors that produced them [...] .”

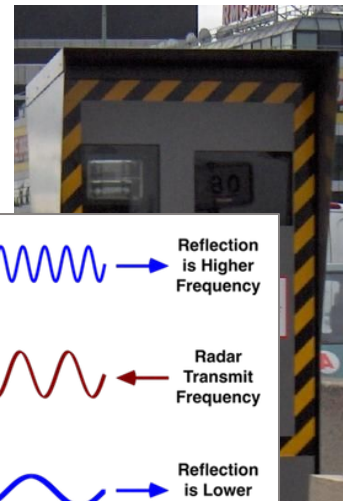
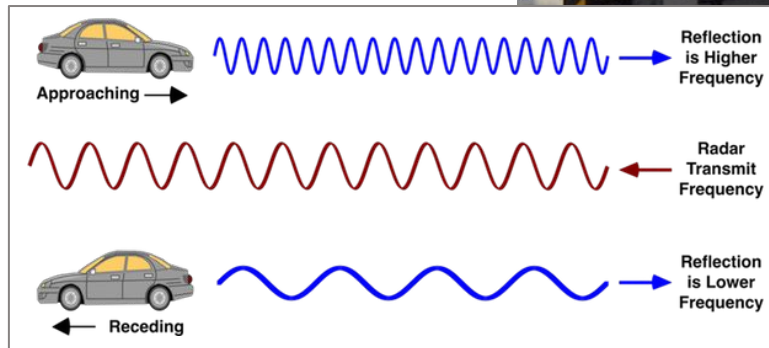
Wikipedia, EN

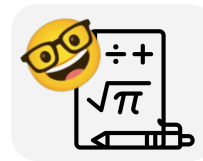


Inverse Problems ?



“Doppler effect”





**Unknown
signal**

x

**Measurement
model**

A



y

**Physical
observations**

- volume
- speed
- image
- time series
- ...

- distance
- frequency
- image
- interferences
- ...

“Forward/design equation”

$$\textcircled{y} = \textcircled{A} \textcircled{x}$$

$$y = Ax$$

“Deconvolution”

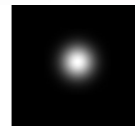
EPFL



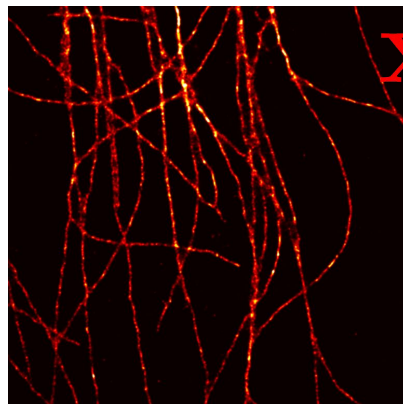
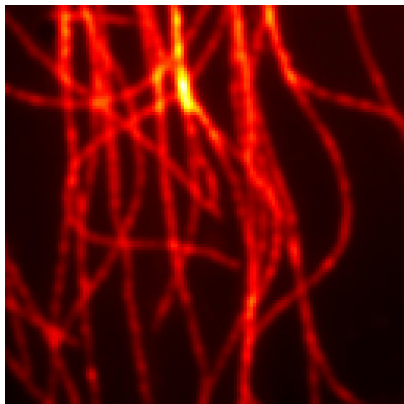
A



A

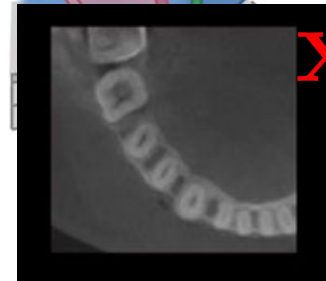
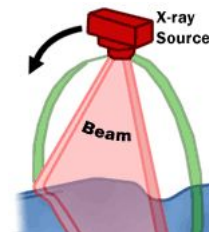
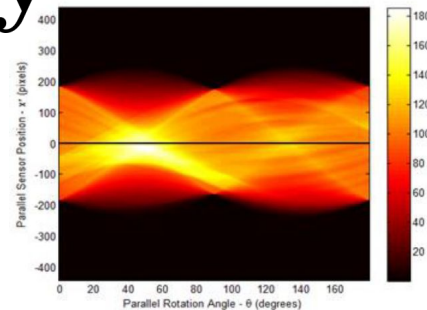


y



x

y



x

“SMLM”

[Credits: Bastien Laville]

[Credits: Wikipedia: CT Scan]

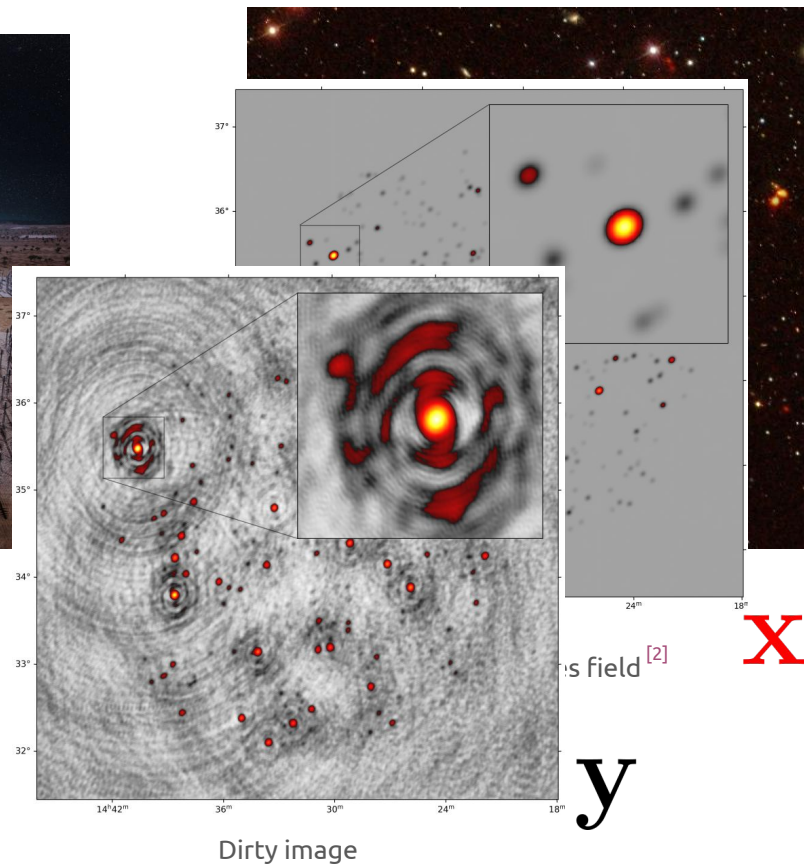
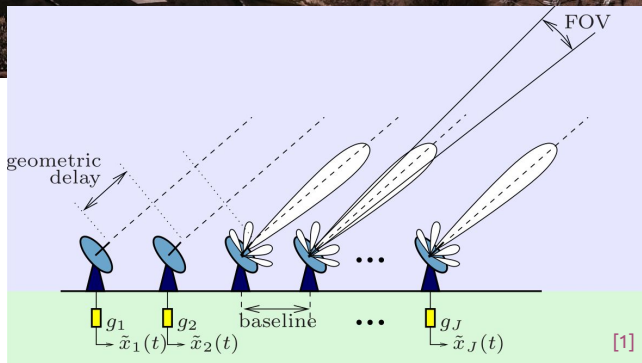
EPFL Radio Interferometric Imaging (still LIP)

Adrian Jarret — Public Defense — 7.11.2025



[Credits: SKAO]

A



[1] Van der Veen et al. "Signal Processing for Radio Astronomy", 2019

[2] Williams WL et al., "LOFAR 150-MHz observations of the Boötes field: catalogue and source counts.", *Monthly Notices of the Royal Astronomical Society*, 2016

$$y = Ax$$

🤔 ~~$x = A^{-1}y$~~

- Measurement noise
- Instability
- Non-uniqueness
- Missing data

$$y = Ax + n$$

$$Ax_1 = Ax_2$$

~~A^{-1}~~



Part I:

What is it about?

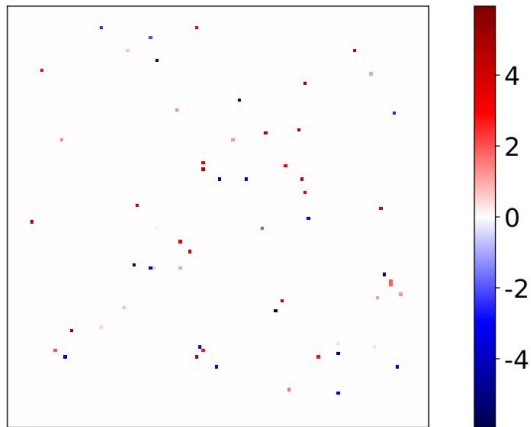
1. ~~Linear Inverse Problems~~
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$$y = Ax$$

“Méthodes atomiques
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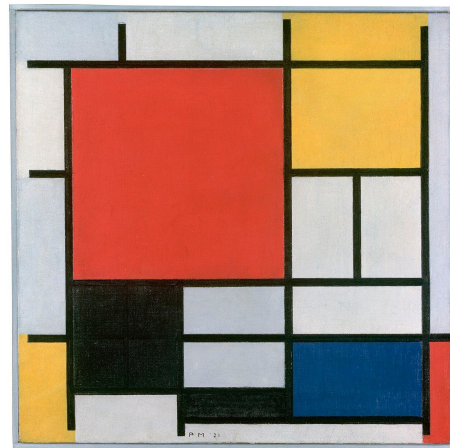
EPFL 2. Parcimonie / Sparsity



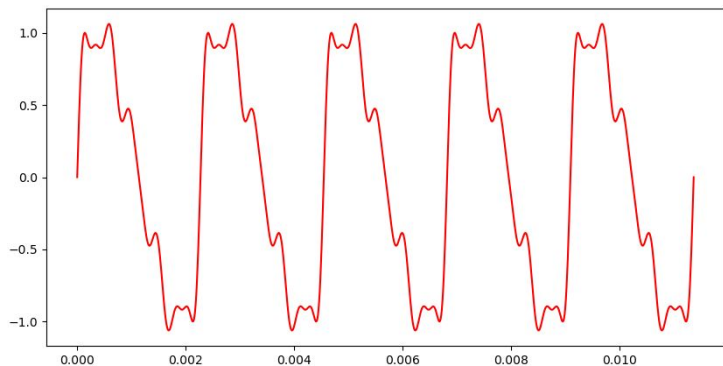
Simulated sparse signal



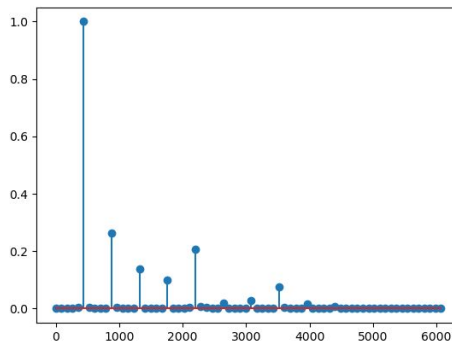
GLEAM observations of the sky



Composition in Red, Yellow, Blue and Black - Piet Mondrian, 1921



A440 violin



$$\mathbf{x} = \alpha_1 \mathbf{a}_1 + \alpha_2 \mathbf{a}_2 + \dots + \alpha_K \mathbf{a}_K$$

EPFL 2. (Bonus) All signals can be approx.-sparse



1% information

EPFL 2. (Bonus) Reconstruction example with sparsity

Something like inpainting or deblurring of an image, with simple sparsity operator

Or Radio Interferometry

Results are maybe not as impressive as what we can see nowadays with AI, but it is controlled, stable and reproducible -> another debate

EPFL 2. (Bonus) Historical moment

Consequence of this approach: change of paradigm, birth of compressed sensing, reduction of the number of observations

Part I:

What is it about?

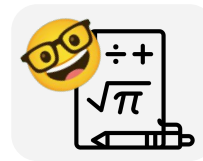
1. ~~Linear Inverse Problems~~
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$$\mathbf{y} = \mathbf{A}\mathbf{x}$$
$$\mathbf{x} = \sum \alpha_k \mathbf{a}_k$$



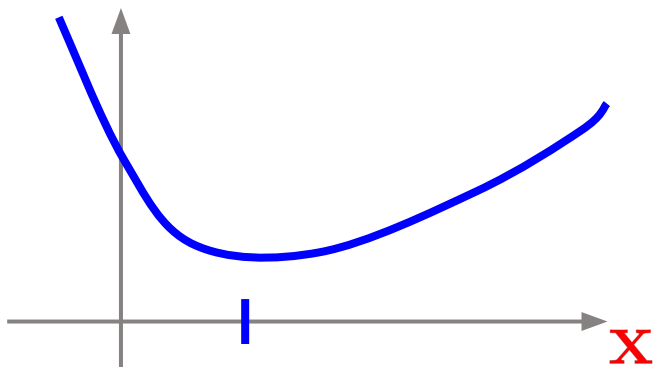
“Méthodes atomiques
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EPFL 3. Penalized Optimization



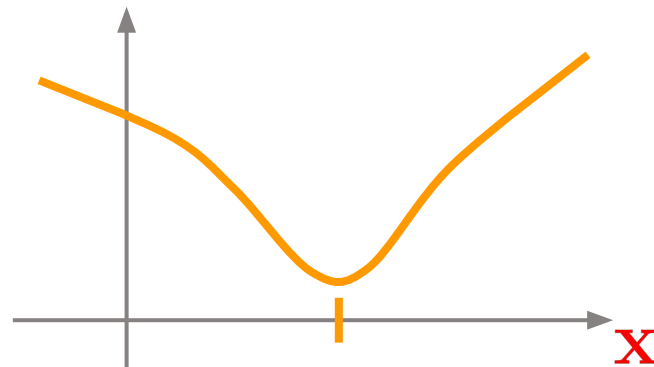
$$\mathbf{y} - \mathbf{A}\mathbf{x} = \mathbf{0}$$

$$E(\mathbf{y} - \mathbf{A}\mathbf{x}) \searrow$$



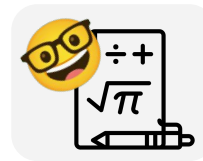
“Data-fidelity”

$$R(\mathbf{x}) \searrow$$

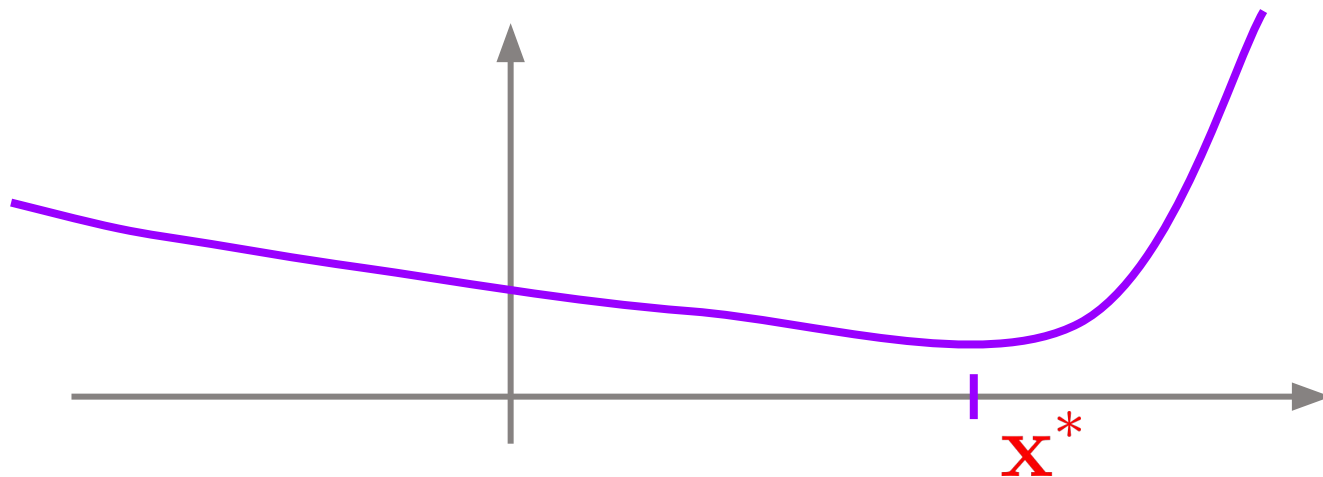


“Penalty/Regularization”

EPFL 3. Penalized Optimization



$$\min_{\mathbf{x}} E(\mathbf{y} - \mathbf{Ax}) + \lambda R(\mathbf{x})$$



LASSO problem^[3] :
$$\min_{\mathbf{x}} \frac{1}{2} \|\mathbf{y} - \mathbf{Ax}\|_2^2 + \lambda \|\mathbf{x}\|_1$$

[3] Tibshirani R. "Regression Shrinkage and Selection via the Lasso", *Journal of the Royal Statistical Society Series B (Methodological)*, 1996.

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What is it about?

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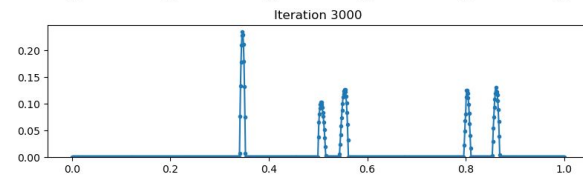
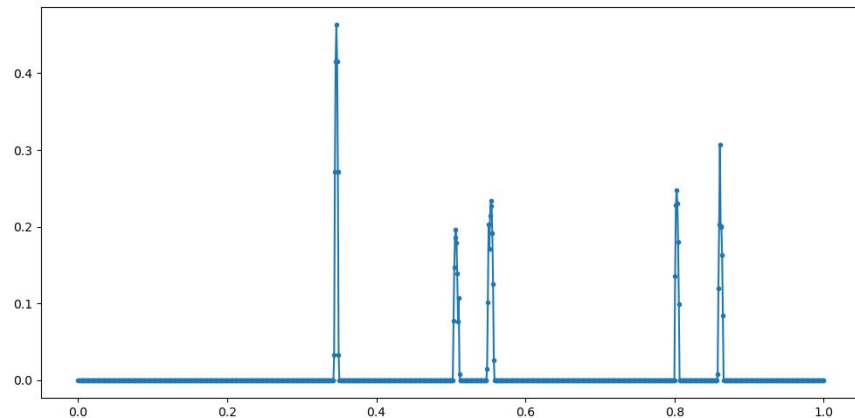
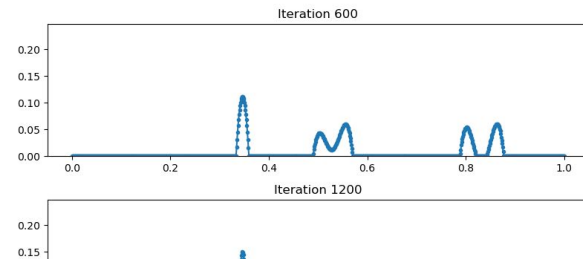
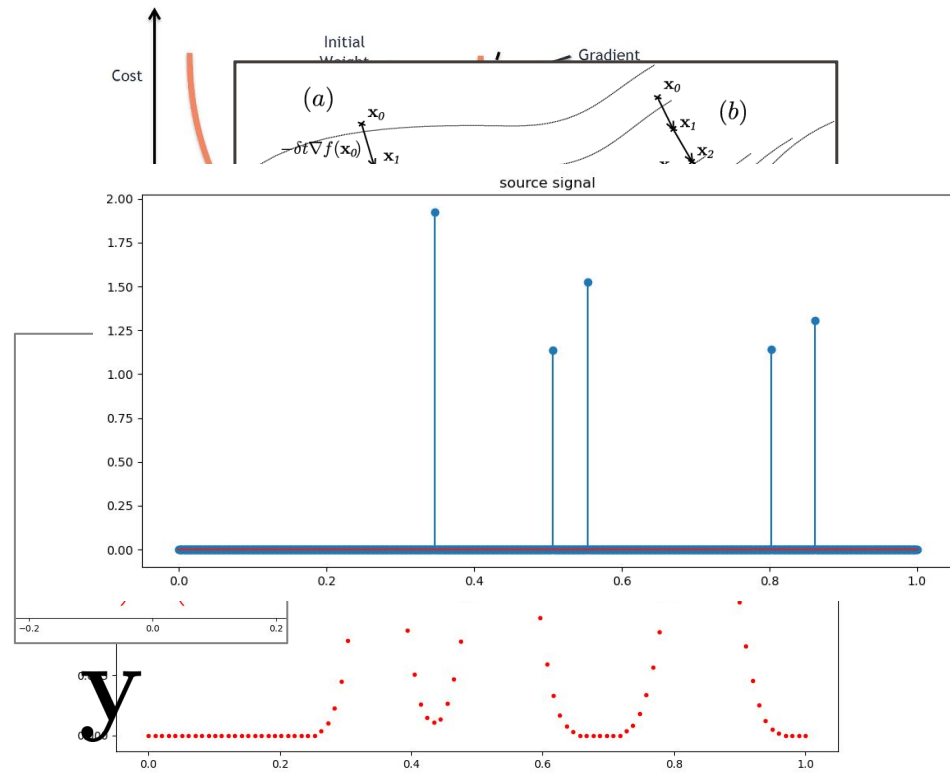
$$\mathbf{y} = \mathbf{A}\mathbf{x}$$
$$\mathbf{x} = \sum \alpha_k \mathbf{a}_k$$
$$\min_{\mathbf{x}} E$$

“**Méthodes atomiques**
pour la **résolution** de
problèmes inverses
parcimonieux”



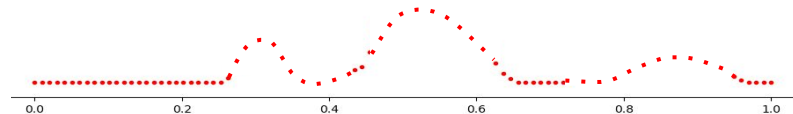
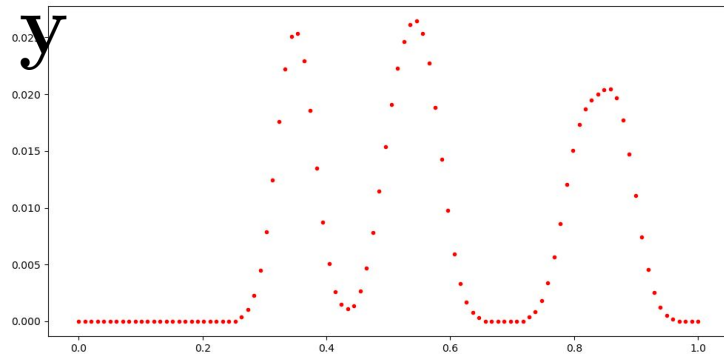
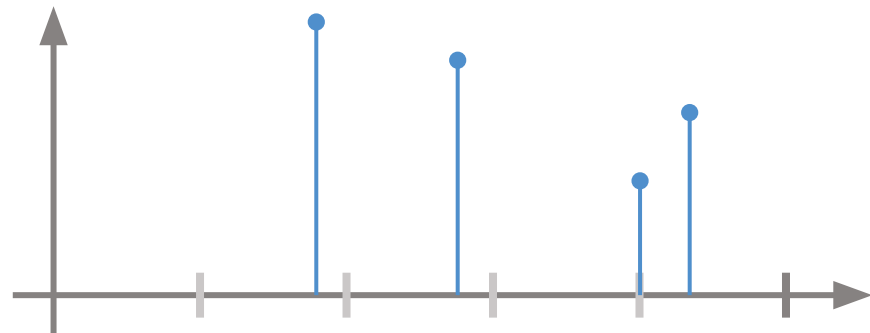
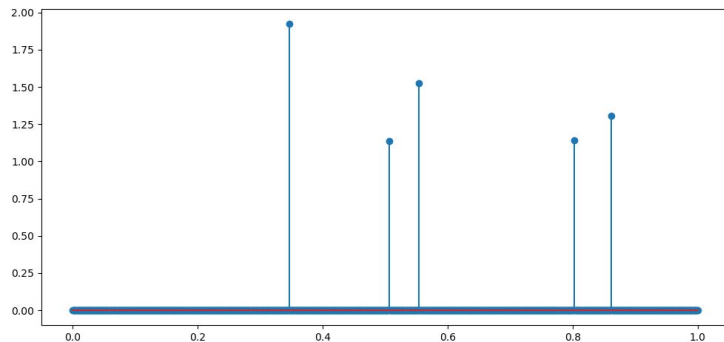
$$\min_{\mathbf{x}} E(\mathbf{y} - \mathbf{A}\mathbf{x}) + \lambda R(\mathbf{x})$$

“Classical” gradient descent:



EPFL 4. Numerical reconstruction methods

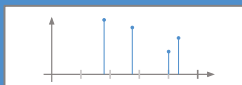
Atomic method:



Part I: What is it about?

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- ~~3. Penalized Optimization~~
- ~~4. Atomic Reconstruction Methods~~
- 5. To Grid or Not To Grid ?**

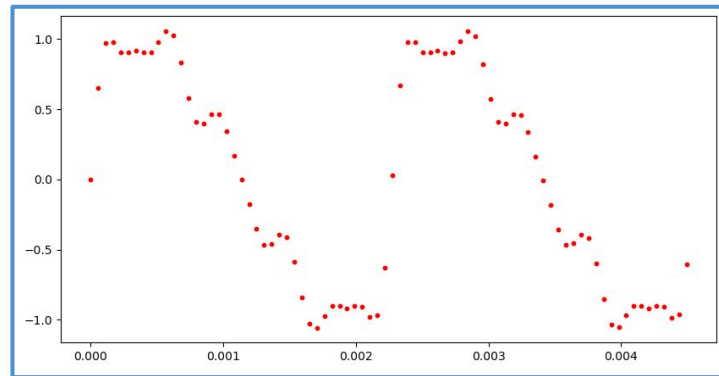
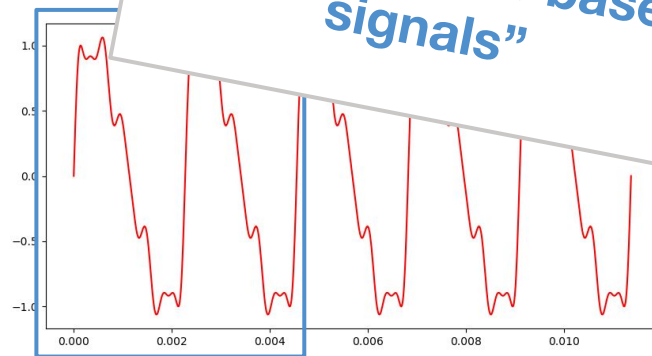
$$\mathbf{y} = \mathbf{A}\mathbf{x}$$
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“Méthodes atomiques
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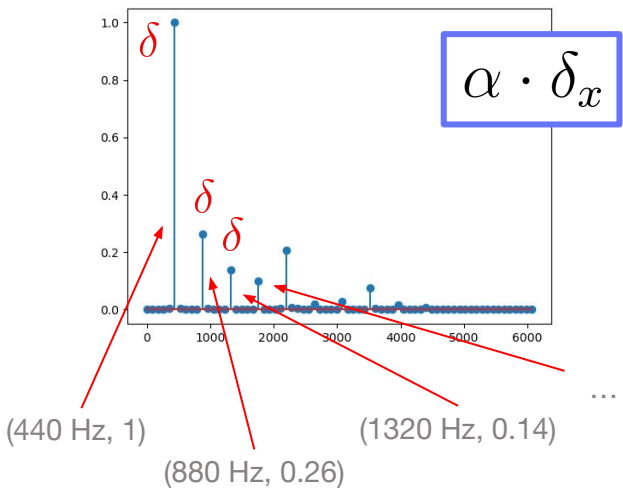
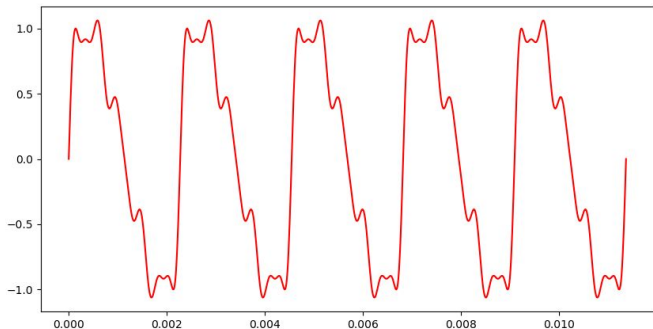
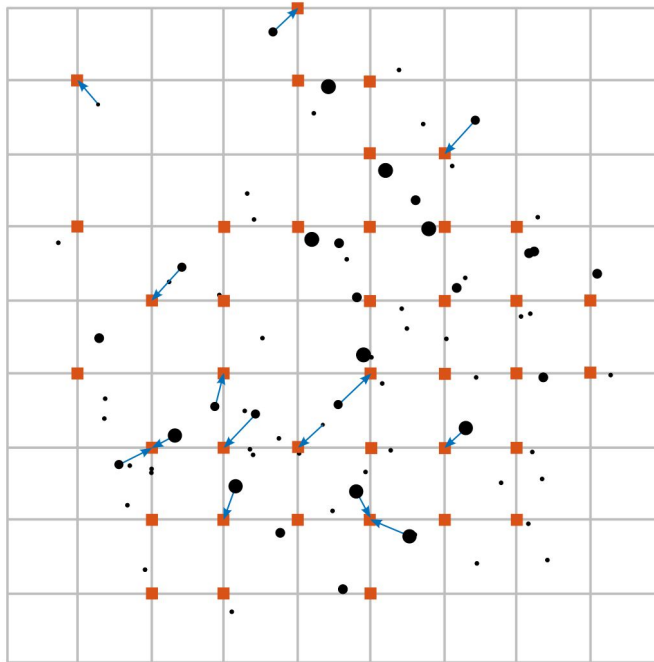
“Discrete/grid-based signals”



EPFL 5. Continuous-domain reconstruction?

$\in \mathcal{M}(\mathbb{R}^d)$

$$\mathbf{x} \in \mathbb{R}^N \quad m = \sum \alpha_k \delta_{x_k}$$

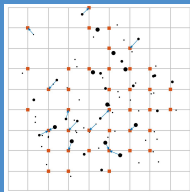
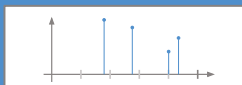


[Credits: H. Pan et al., "LEAP", A&A, 2017.]

Part I: What is it about?

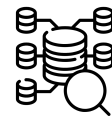
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$$\mathbf{y} = \mathbf{A}\mathbf{x}$$
$$\mathbf{x} = \sum \alpha_k \mathbf{a}_k$$
$$\min_{\mathbf{x}} E$$



Challenges:

- Time efficiency
- Accuracy
- Scalability



+ Radio Interferometry

Part II: Focus on key contributions

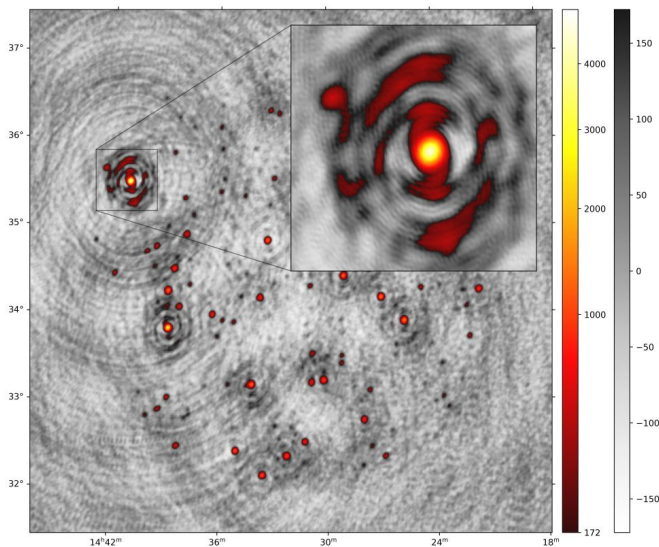
1. **PolyCLEAN:
A Polyatomic Method for
Radio Interferometry**
2. **Decoupling of Composite
Problems**

EPFL Radio Interferometric Imaging - Dirty Image

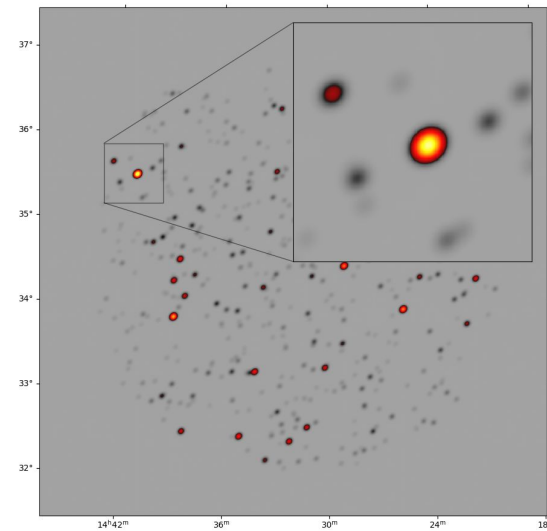
$$\mathbf{V} = \Phi \mathbf{I} + \epsilon$$

$$\mathbf{I}_D = \Phi^* \mathbf{V}$$

$\mathbf{I}$



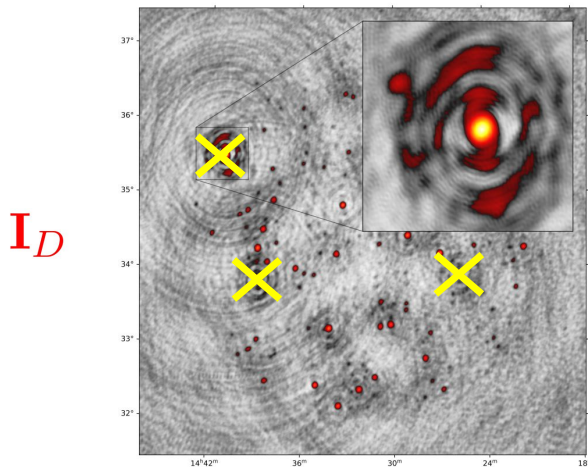
Dirty image from observations



Catalogue Boötes field^[4]

[4] Williams WL et al., "LOFAR 150-MHz observations of the Boötes field: catalogue and source counts.", *Monthly Notices of the Royal Astronomical Society*, 2016

The CLEAN family^[3]



Intuitive and simple method,
long-developed and fast



Sensitive to stop, objective function
unclear, physically impossible artefacts

Optimization-based Methods^[4]

$$\arg \min_{\mathbf{I} \in \mathbb{R}^N} \frac{1}{2} \|\mathbf{y} - \Phi \mathbf{I}\|_2^2 + \mathcal{R}(\mathbf{I})$$

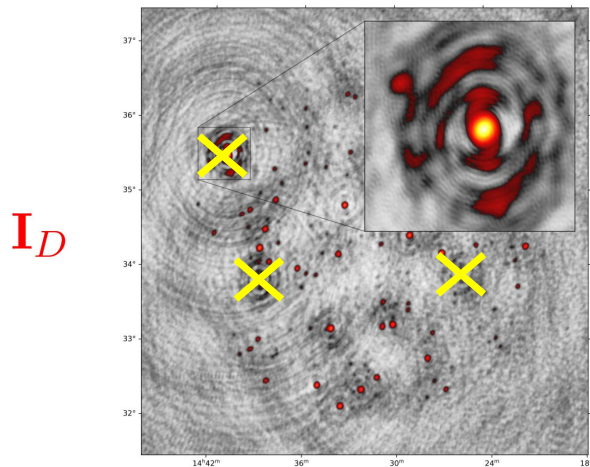


Controlled solutions, versatile priors,
excellent results, additional tools

[3] Högbom JA., "Aperture Synthesis with a Non-Regular Distribution of Interferometer Baselines", *Astronomy and Astrophysics Supplement Series*, 1974.

[4] Wiaux Y. et al., "Compressed sensing imaging techniques for radio interferometry", *Monthly Notices of the Royal Astronomical Society*, 2009.

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LASSO



Controlled solutions, versatile priors,
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Numerically heavy, little adoption in the
field

[3] Högbom JA., "Aperture Synthesis with a Non-Regular Distribution of Interferometer Baselines", *Astronomy and Astrophysics Supplement Series*, 1974.

[4] Wiaux Y. et al., "Compressed sensing imaging techniques for radio interferometry", *Monthly Notices of the Royal Astronomical Society*, 2009.

EPFL Our Polyatomic Frank-Wolfe Algorithm

$$\arg \min_{\mathbf{x} \in \mathbb{R}^N} \mathcal{J}(\mathbf{x}) := \frac{1}{2} \|\mathbf{y} - \mathbf{A}\mathbf{x}\|_2^2 + \lambda \|\mathbf{x}\|_1$$

Algorithm 3: Polyatomic Frank-Wolfe algorithm of quality δ ^[9]

Initialize $\mathbf{x}_0 \in \mathcal{D}$, $\mathcal{S}_0 = \emptyset$

for $k = 1, 2 \dots$ **do**

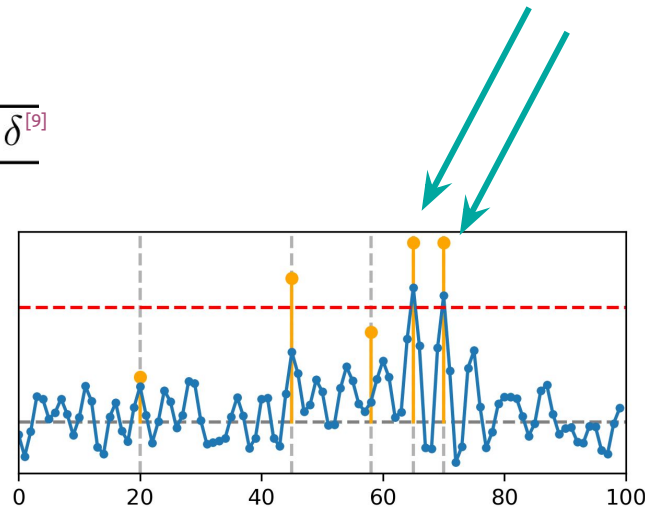
1) Find update directions:

$$\mathcal{I}_k \leftarrow \{1 \leq j \leq N : |\boldsymbol{\eta}_k[j]| \geq \|\boldsymbol{\eta}\|_\infty - \delta/k\}$$

$$\mathcal{S}_k \leftarrow \mathcal{S}_{k-1} \cup \mathcal{I}_k$$

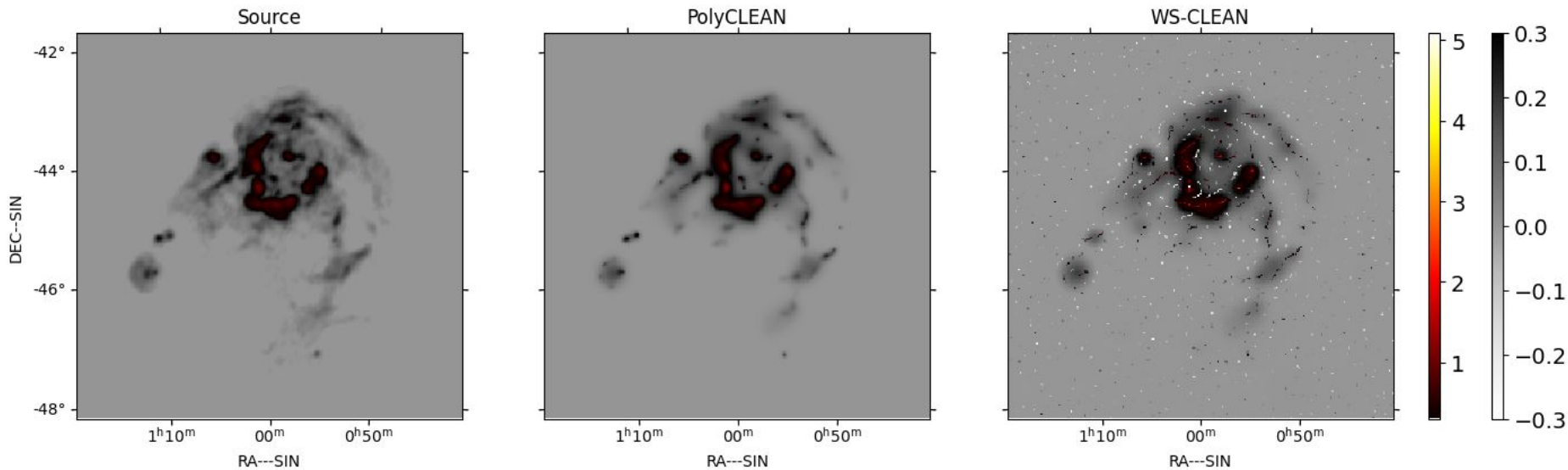
2) Reweight:

$$\mathbf{x}_k \leftarrow \arg \min_{\text{Supp}(\mathbf{x}) \subset \mathcal{S}_k} \frac{1}{2} \|\mathbf{y} - \mathbf{A}\mathbf{x}\|_2^2 + \lambda \|\mathbf{x}\|_1$$

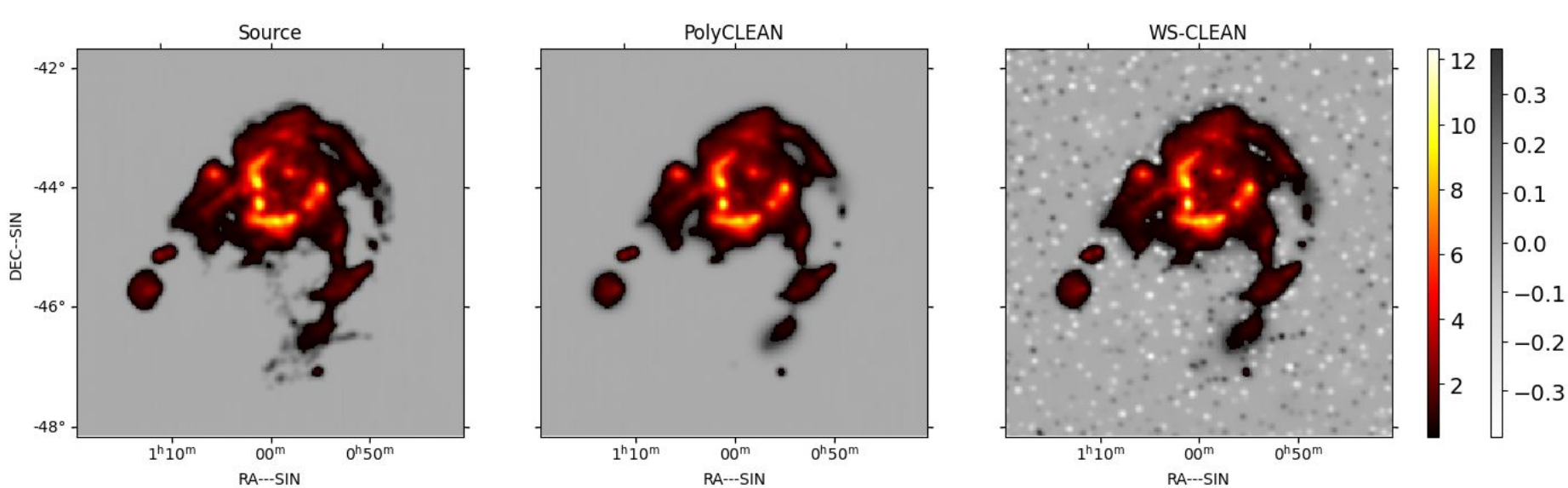


$$|\mathcal{S}_k| \leq N$$

EPFL Diffuse emission reconstruction



EPFL Diffuse emission reconstruction





Design of optimization algorithm → Real world application



Best of both worlds:

- Benefits of CLEAN → Atomic, fast
- Benefits of convex optimization → Accurate
- Sparsity-aware processing (HVOX) → Numerical efficiency
- Question of resolution
→ CLEAN beam is too coarse, certificate beam is data-inspired

Part II: Focus on key contributions

- ~~1. PolyCLEAN:
A Polyatomic Method for
Radio Interferometry~~
2. Decoupling of Composite Problems

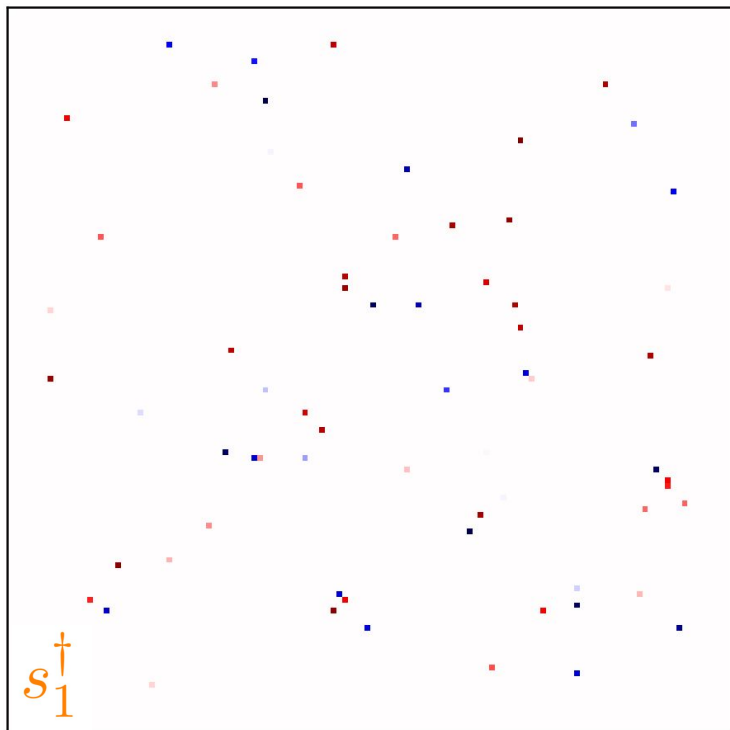
EPFL Sparse-plus-Smooth Problems



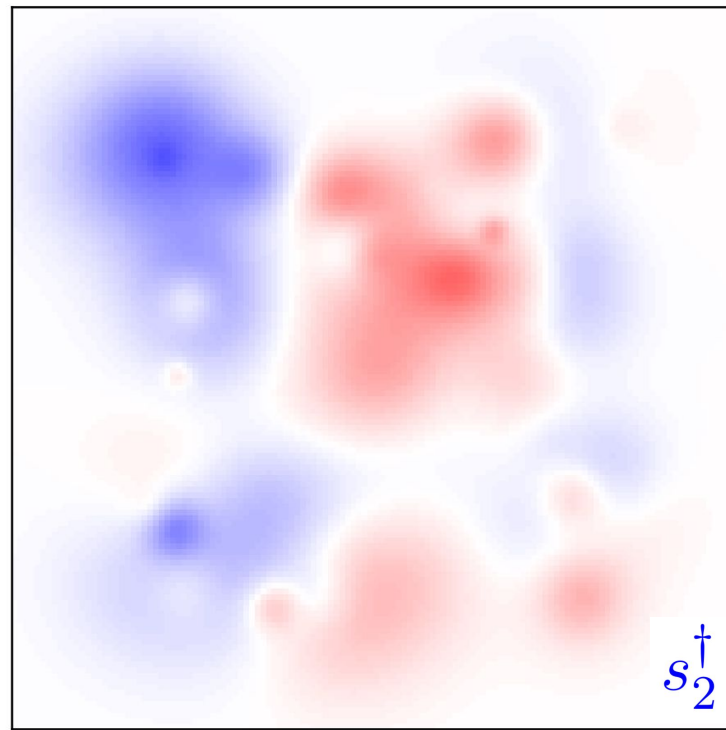
GLEAM survey of the radio sky, J2000 coordinates (9h37min15.21s, 50°25'03.1")

EPFL Sparse-plus-Smooth Problems

$$s_1^\dagger + s_2^\dagger$$



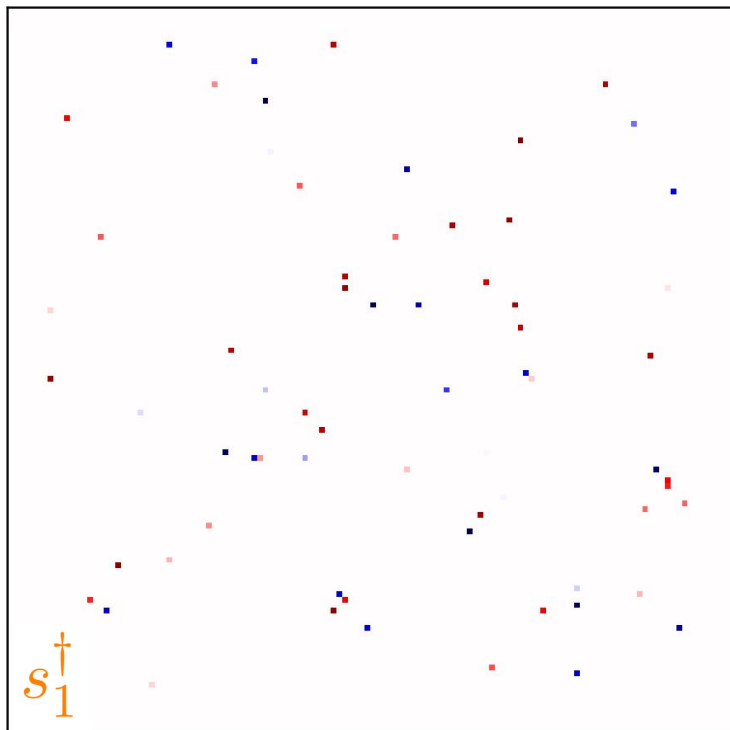
Sparse foreground



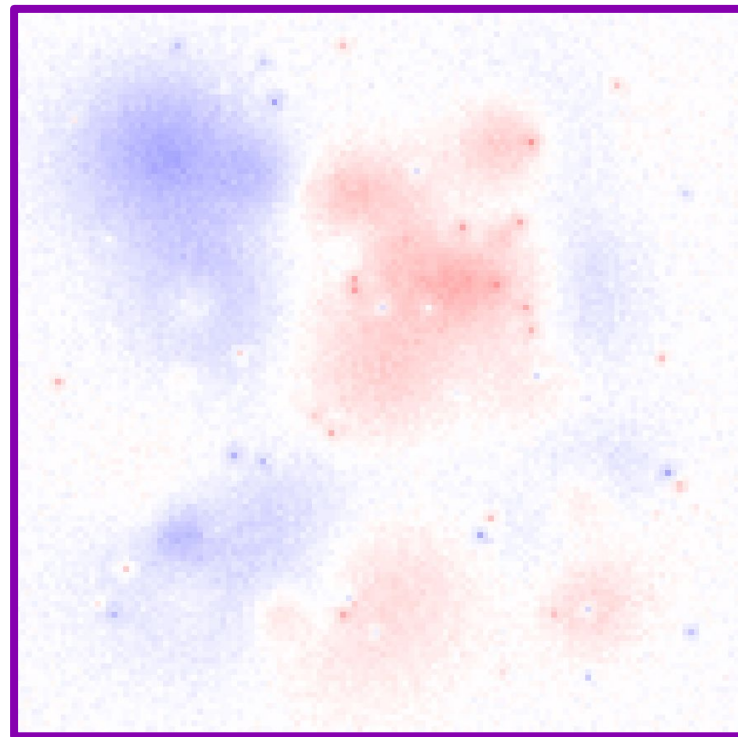
Smooth background

EPFL Sparse-plus-Smooth Problems

Adrian Jarret — Public Defense — 7.11.2025



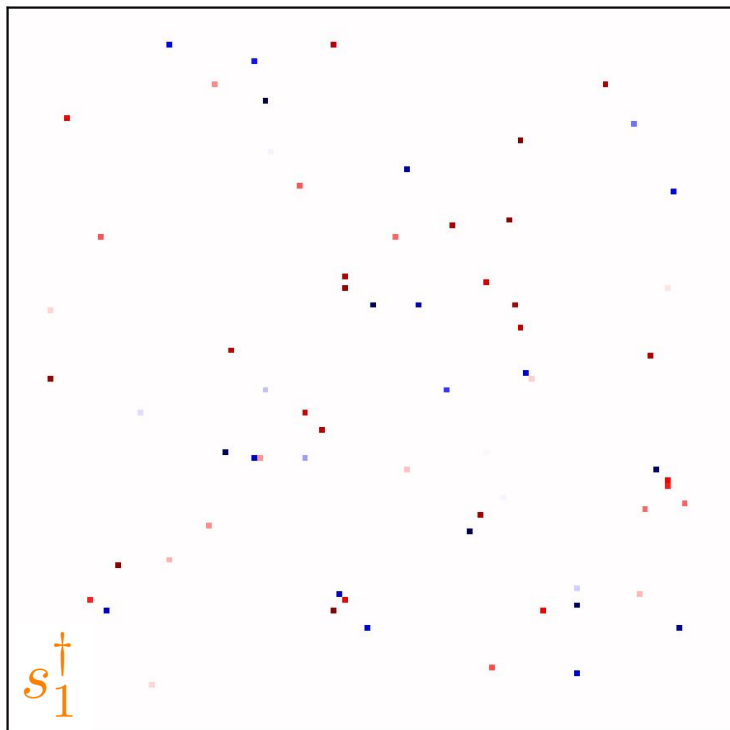
Sparse foreground



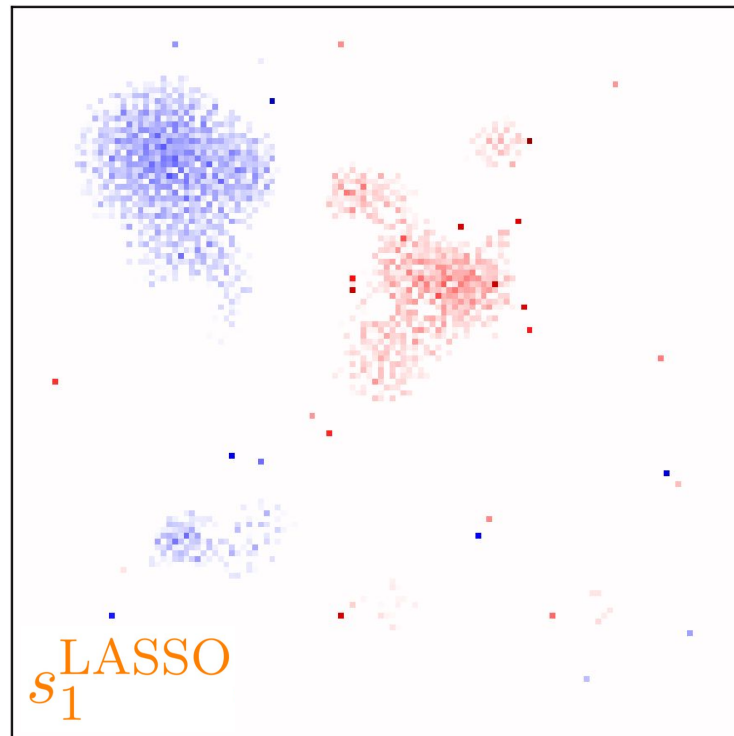
Dirty Image

EPFL Sparse-plus-Smooth Problems

Adrian Jarret — Public Defense — 7.11.2025

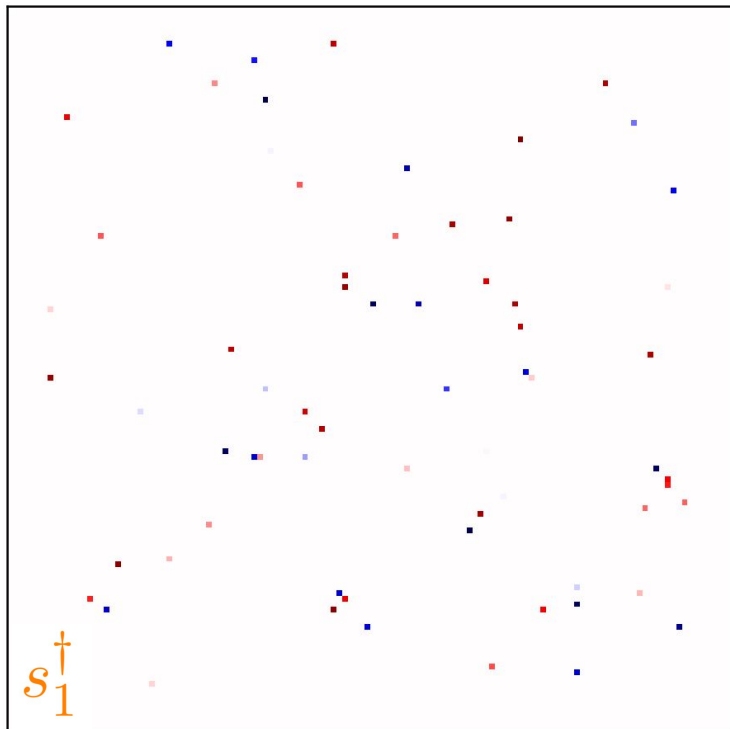


Sparse foreground

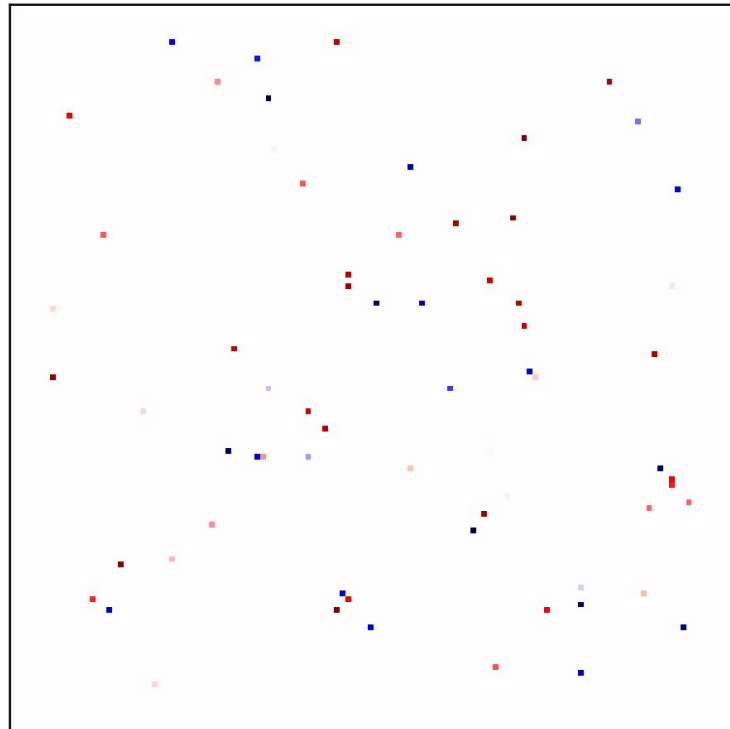


LASSO reconstruction

$$\arg \min_{s_1, s_2} \mathcal{J}(s_1, s_2)$$



Sparse foreground



Composite model

EPFL Our Composite Representer Theorem

$$\arg \min_{\mathbf{s}_1, \mathbf{s}_2 \in \mathcal{M}(\mathcal{X}) \times \mathbf{L}_2(\mathcal{X})} \frac{1}{2} \|\mathbf{y} - \Phi(\mathbf{s}_1 + \mathbf{s}_2)\|_2^2 + \lambda_1 \|\mathbf{s}_1\|_{\mathcal{M}} + \frac{\lambda_2}{2} \|\mathbf{s}_2\|_{\mathbf{L}_2}^2$$

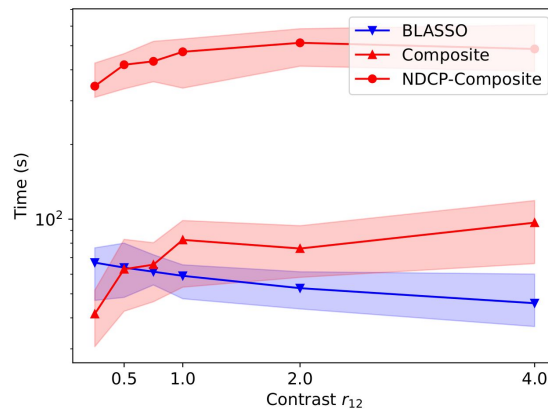
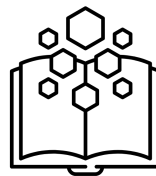
Our Representer Theorem

[Theorem 6.1]

$$\begin{cases} \hat{\mathbf{s}}_1 \in \arg \min_{\mathbf{s}_1 \in \mathcal{B}} \frac{1}{2} \|\mathbf{M}_{\lambda_2}^{-\frac{1}{2}} (\mathbf{y} - \Phi(\mathbf{s}_1))\|_2^2 + \lambda_1 \|\mathbf{s}_1\|_{\mathcal{B}} \\ \hat{\mathbf{s}}_2 = \frac{1}{\lambda_2} \Phi^* \mathbf{M}_{\lambda_2}^{-1} (\mathbf{y} - \mathbf{w}) \end{cases}$$

$$\mathbf{M}_{\lambda_2} := \frac{1}{\lambda_2} (\Phi \Phi^* + \lambda_2 \mathbf{I}_L)$$
$$\mathbf{w} = \Phi(\hat{\mathbf{s}}_1)$$

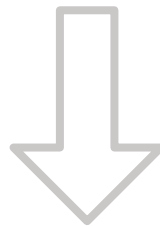
- Better understanding of the composite optimization problem
- Nice mathematical framework
- More efficient optimization algorithms



A Short Summary of this Thesis

1. Polyatomic Frank-Wolfe
2. PolyCLEAN
3. Continuous-Domain PFW
4. Composite Reconstruction

**Better understanding
of the
problems**



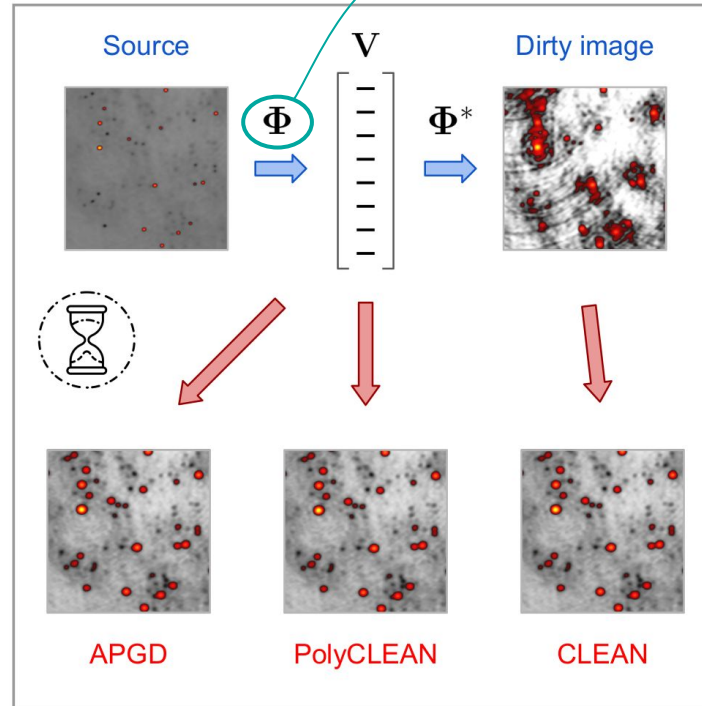
**Develop more
efficient
numerical methods**



Thank you !

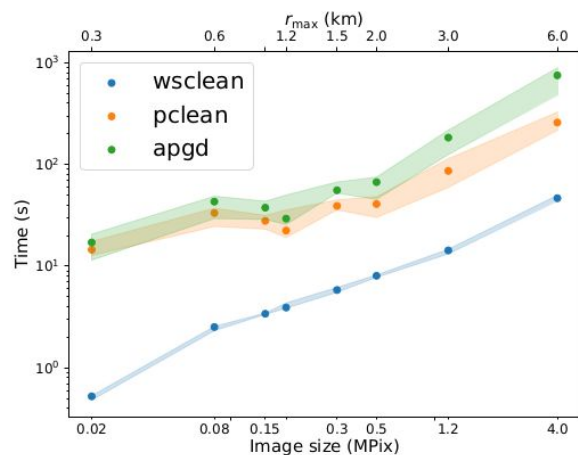
Supplementary slides

A Fast Reconstruction Method

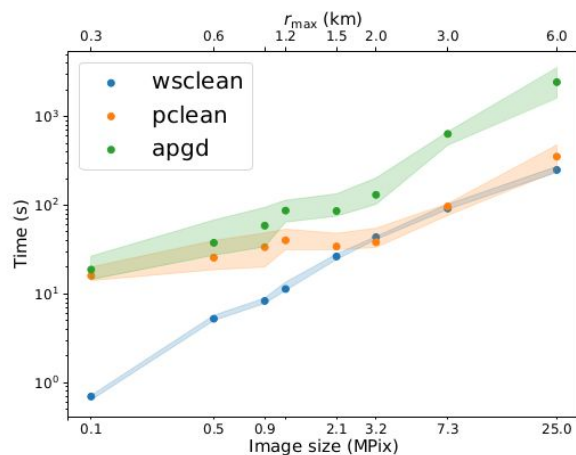


Varying number of
baselines

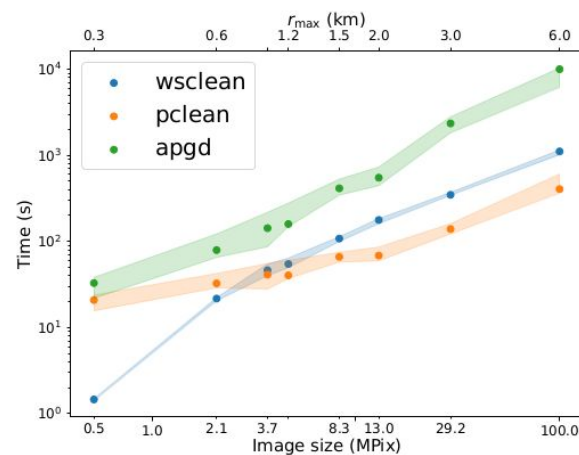
A Fast Reconstruction Method



SRF 2



SRF 5



SRF 10

[11] Jarret A et al., "PolyCLEAN: Atomic optimization for super-resolution imaging and uncertainty estimation in radio interferometry", *Astronomy & Astrophysics*, 2025.

Benefits of Polyatomic Frank-Wolfe

- Polyatomic $\rightarrow$ **Fast**
- Sparse iterates $\rightarrow$ **Scalable**
- Convergence $\rightarrow$ **Principled**

$$\mathcal{S}_k \leftarrow \mathcal{S}_{k-1} \cup \mathcal{I}_k$$

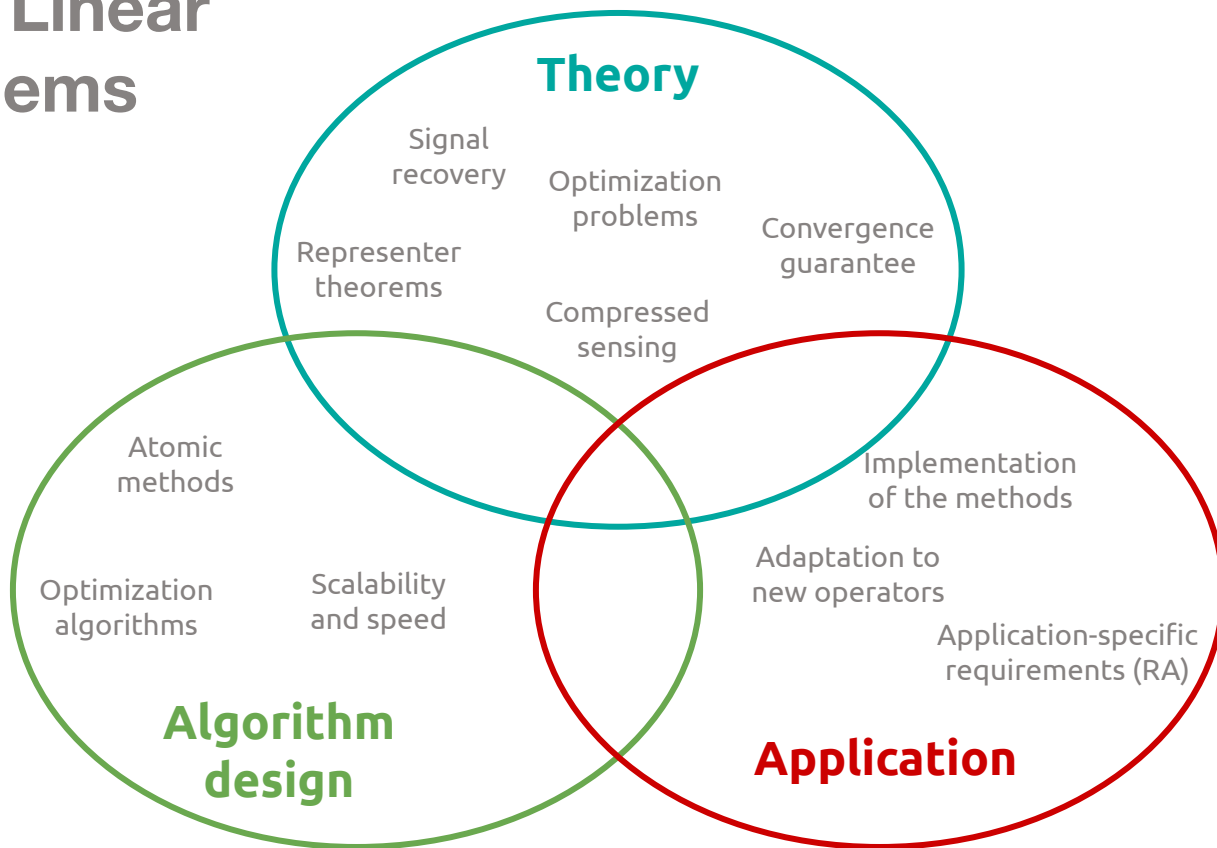
$$\mathbf{x}_k = \sum_{i=1}^{N_k} \alpha_i^{[k]} \mathbf{e}_i^{[k]}$$

Theorem 3.2 (Convergence of Polyatomic FW).^[9]

$$\mathcal{J}(\mathbf{x}_k) - \mathcal{J}^* \leq \frac{2}{k+2} (C_f + 2\delta)$$

[9] Jarret A, Fageot J, Simeoni M. "A Fast and Scalable Polyatomic Frank-Wolfe Algorithm for the LASSO", *IEEE Signal Processing Letters*, 2022.

Landscape of **Linear Inverse Problems**



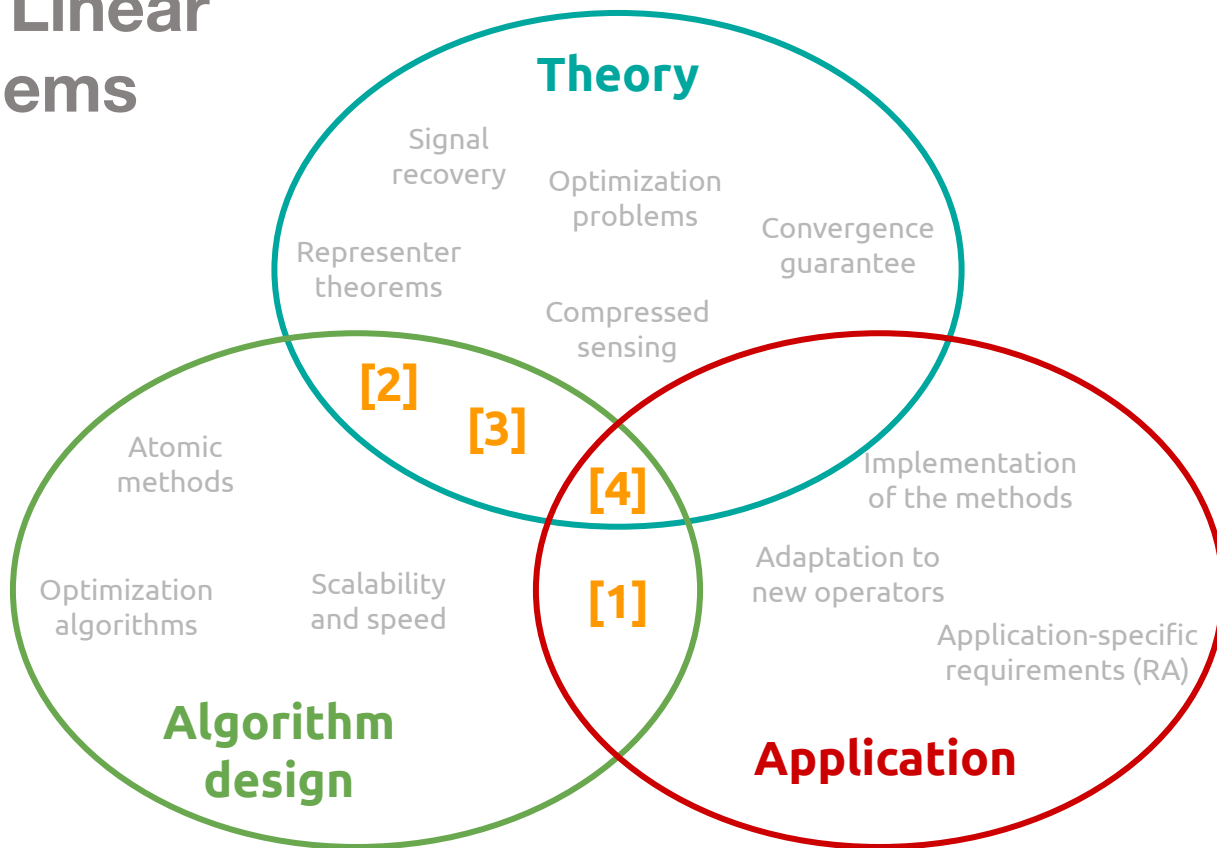
Landscape of Linear Inverse Problems

[1] PFW

[2] CD-PFW

[3] Composite

[4] PolyCLEAN



Conclusion and perspectives

Mathematics-aware numerical solvers:

- Principled (poly)atomic methods
- Decoupled algorithms
- Sparsity-aware processing

Chapters 3, 5

Chapter 6

Chapter 8

Open questions:

- Resolution of the reconstruction and quantitative imaging
- More advanced *traditional* methods
- Mixed learning-based approaches

Composite Representer Theorem (in the literature)

$$\arg \min_{s_1, s_2 \in \mathcal{M}(\mathcal{X}) \times L_2(\mathcal{X})} \frac{1}{2} \|\mathbf{y} - \Phi(s_1 + s_2)\|_2^2 + \lambda_1 \|s_1\|_{\mathcal{M}} + \frac{\lambda_2}{2} \|s_2\|_{L_2}^2 \quad [13]$$

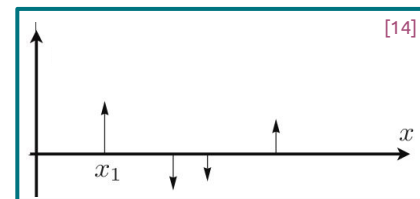
Representer theorem^[13]:

- $s_1^* \rightarrow$ Sparse measure

$$s_1^* = m[\mathbf{a}^*, \mathbf{x}^*]$$

- $s_2^* \rightarrow$ Smooth quadratic solution

$$s_2^* = \Phi^* \mathbf{u}^*$$



[13] Debarre T et al. "Continuous-Domain Formulation of Inverse Problems for Composite Sparse-Plus-Smooth Signals", *IEEE Open Journal of Signal Processing*, 2021.

[14] Unser M et al., "Splines Are Universal Solutions of Linear Inverse Problems with Generalized TV Regularization", *SIAM Review*, 2017.